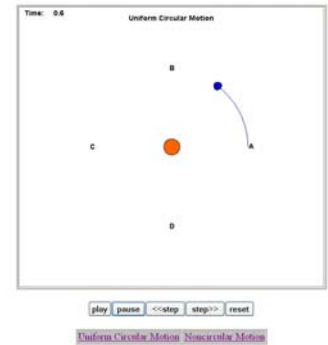


RECITATION HANDOUT #13: Gravitational Orbits

Properties of Circular vs. Elliptical Orbits

There are **two types** of closed orbits for a planetary system: **circular** and **elliptical** (circle = special ellipse). Use http://www.courses.vcu.edu/PHYS-jms/PHYS207/gravity/circ_ellip_motion.htm to answer the following questions about orbital motion for these two cases. The inner circle represents the sun (mass M) and the outer circle represents a planet (mass m) orbiting the sun. *We assume that the sun remains fixed in place.*



Circular Orbit

- (a) In the circular orbit, **how long** does it take the planet to go from point A to point B? _____ s
- (b) **How long** does it take the planet to go from point B to point C? _____ s
- (c) At **what point** (A,B,C,D, or all the same) is the **kinetic energy** maximum? _____
- (d) At **what point** (A,B,C,D, or all the same) is the **potential energy** maximum? _____
Remember that the gravitational potential energy equals $-\frac{GMm}{r}$.
- (e) How do the **gravitational force** $\frac{GMm}{r^2}$ and **centripetal force** $\frac{mv^2}{r}$ compare at points A and C?

(circle the correct relationship)

Point A: $\frac{GMm}{r_A^2} > \frac{mv^2}{r_A}$, $\frac{GMm}{r_A^2} < \frac{mv^2}{r_A}$, or $\frac{GMm}{r_A^2} = \frac{mv^2}{r_A}$

Point C: $\frac{GMm}{r_C^2} > \frac{mv^2}{r_C}$, $\frac{GMm}{r_C^2} < \frac{mv^2}{r_C}$, or $\frac{GMm}{r_C^2} = \frac{mv^2}{r_C}$

Elliptical Orbit

- (a) In the elliptical orbit, **how long** does it take the planet to go from point A to point B? _____ s
- (b) **How long** does it take the planet to go from point B to point C? _____ s
- (c) At **what point** (A,B,C,D, or all the same) is the **kinetic energy** maximum? _____
- (d) At **what point** (A,B,C,D, or all the same) is the **potential energy** maximum? _____
Remember that the potential energy $-\frac{GMm}{r}$ increases (or becomes less negative) as the radius increases. Also, no external force acts on the system, so total energy is conserved.
- (e) How do the **gravitational force** $\frac{GMm}{r^2}$ and **centripetal force** $\frac{mv^2}{r}$ compare at points A and C?

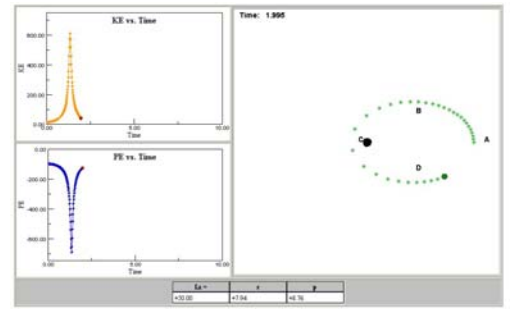
Point A: $\frac{GMm}{r_A^2} > \frac{mv^2}{r_A}$, $\frac{GMm}{r_A^2} < \frac{mv^2}{r_A}$, or $\frac{GMm}{r_A^2} = \frac{mv^2}{r_A}$

Point C: $\frac{GMm}{r_C^2} > \frac{mv^2}{r_C}$, $\frac{GMm}{r_C^2} < \frac{mv^2}{r_C}$, or $\frac{GMm}{r_C^2} = \frac{mv^2}{r_C}$

Angular Momentum and Energy of Elliptical Orbits

Use http://www.courses.vcu.edu/PHYS-jms/PHYS207/gravity/ke_pe_orbit.htm

to answer the following questions. Assume that the mass of the “planet” is **2 kg** and that the mass of the sun is such that $GM_{sun} = 1000$. We also assume that the sun remains fixed in place.



(a) With the initial conditions such that $r_o = 10$ m and $v_{oy} = 5$ m/s, what **type of orbit** does the planet make? (circular, elliptical, hyperbolic, parabolic)

(b) What is the **angular momentum** of the planet? Hint: L_z is given on the applet page.

$$L_z = \boxed{} \text{ kg m}^2/\text{s}$$

(c) Is there an **x or y component** of the angular momentum? Why or why not?

(d) In general, does $L_z = r p$, where r is the distance of the planet to the sun, and p is the magnitude of the momentum? Why or why not?

(e) Is the angular momentum of the planet **constant** during the orbit? Why or why not?

(f) The **kinetic** energy is **maximum** at point _____ and **minimum** at point _____.

(g) The **potential** energy is **maximum** at point _____ and **minimum** at point _____.

Conditions for Circular Orbits

(h) Can you create a stable **circular** orbit? With $r_o = 10$ m, what **velocity** v_{y0} is required to produce a circular orbit? Try different values for v_{y0} in the physlet. Can you justify your answer using the relationship between the gravitational and centripetal forces for a circular orbit?

$$\text{For } r_o = 10 \text{ m, } v_{y0} = \boxed{} \text{ m/s}$$

(i) If instead you start with $r_o = 5$ m, what **velocity** v_{y0} is required to produce a circular orbit? How does this velocity compare with the larger radius in part (h)?

$$\text{For } r_o = 5 \text{ m, } v_{y0} = \boxed{} \text{ m/s}$$

(j) If instead you start with $r_o = 40$ m, what **velocity** v_{y0} is required to produce a circular orbit? You will need to calculate this value – look at the hint given in part (h).

$$\text{For } r_o = 40 \text{ m, } v_{y0} = \boxed{} \text{ m/s}$$